

Air Force Research Laboratory



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Intercalibration

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- Overview
- Conjunction criteria
- Preparing data sets
- Calculation of statistics
- Sample results for electrons
- Results for data cleaning
- Summary







- On-orbit intercalibration of instruments is required to—
 - Determine systematic offsets between data sets (bias)
 - Determine measurement uncertainty (random error)
- Data sets are corrected for bias relative to an instrument nominally identified as a "gold" standard
 - Proton standard sensor: GOES 8/SEM
 - Electron standard sensor: CRRES/MEA+HEEF
- Random error is an input into development of flux maps
- Intercalibration is based on comparisons of joint observations
 - easier for protons—use simultaneous SPE observations
 - harder for electrons—focus of this talk



Proton CrossCal Tree





Electron CrossCal Tree





Conjunction Criteria



- For satellites beyond LEO our conjunction criteria is patterned after that in Friedel et al., 2005, *Space Weather*, 3:S09B04
 - Match L*, B/Bo, and UT
 - Assume little MLT variation
 - Restrict to quiet times
 - Adjust constraints to get necessary number of conjunctions
- For LEO conjunctions, we include some geographic constraints (e.g. GLON)
 - K-h_{min} criteria is another option

Criterion	AE9 CRRESGEO	Friedel et al.
L*	<6.5	<6.0
ΔL*	<0.1	<0.1
Δ(B/Bo)	<0.1	<0.1
ΔUT	<3-4 hr	<3 hr
MLT	4-8 or 16-20	4-8 or 16-20
ΔΜΙΤ	N/A or <2 hr	<2 hr
Кр	<3 last 48 hr	<2 last 48 hr

Criterion	AE9 TSX-5SAMPEX
L*	2.5 <l*<6.5< td=""></l*<6.5<>
ΔL*	<(lesser of 0.1 and 5%)
Δ(B/Bo)	<0.1
ΔUT	<4 hr
ΔGLON	<60 deg
GLAT	same hemisphere







- Intercalibration is done with the lowest level of data possible
 - For example, dosimeter channel results rather than results from inversions
- Comparisons generally require constructing comparable channels
 - Interpolate from standard channel energies to energies of target instrument
 - Integrate from standard differential channels to compare to target integral channels
- In general omnidirectional averaged data is used







The principal intercalibration statistics needed for incorporating data sets are referred to as *cbias* and *dlnj*.

- Take J_A and J_B, series of joint flux observations by satellites A and B, respectively (A=standard, B=target), for a single energy channel.
- Find median ratio

 $R = median(\mathbf{J}_{B} / \mathbf{J}_{A}).$

- The bias of J_B relative to J_A is described by *cbias* = ln(R).
- The satellite B series is adjusted

 $\mathbf{J'}_{\mathrm{B}} = \mathbf{J}_{\mathrm{B}}/\mathrm{R},$

so that the series J_A and J'_B have the same medians.

• The residual error is

 $\mathbf{RE}=\ln(\mathbf{J'}_{\mathsf{B}}/\mathbf{J}_{\mathsf{A}}).$

• The random error of series J_B is $dlnj=[(1/n)(\Sigma RE^2)]^{0.5}$.





Results for CRRES-GEO (1)



9







Comparison of GEO 1989-046 SOPA to CRRES MEA+HEEF









- Intercalibration results should be checked for information on issues such as data cleaning
- Plot shows intercalibration for POLAR HISTe vs. GPS NS24 BDDII, >1.12 MeV electrons
 - Roll-over of POLAR fluxes consistent with paralyzing deadtime issue in HISTe
 - Noise floor in POLAR too
 - Standard data cleaning process will show these issues, but deadtime issues in particular are more obvious here
 - Reminder—the more information on the instrument, the better









- The main objective of intercalibration is to obtain *cbias* and *dlnj* values
- Ideally, we seek to be able to trace a calibration chain back to a standard instrument
- Conjunction/event criteria are adjusted based on frequency of events
- We can work with you on intercalibration, or do this for your data set given enough information on your instrument





Questions & Discussion



